1st ELSI International Symposium

Origin of Earth's Ocean Hidenori Genda

(Univ. Tokyo → ELSI)

- How much water is on the Earth?
- When did the ocean form?
- Where did the water come from?

Earth's Water

How much water?

Ocean: 1.4 x10²¹kg

Earth : $6.0 \times 10^{24} \text{kg}$

tiny fraction (0.023%)

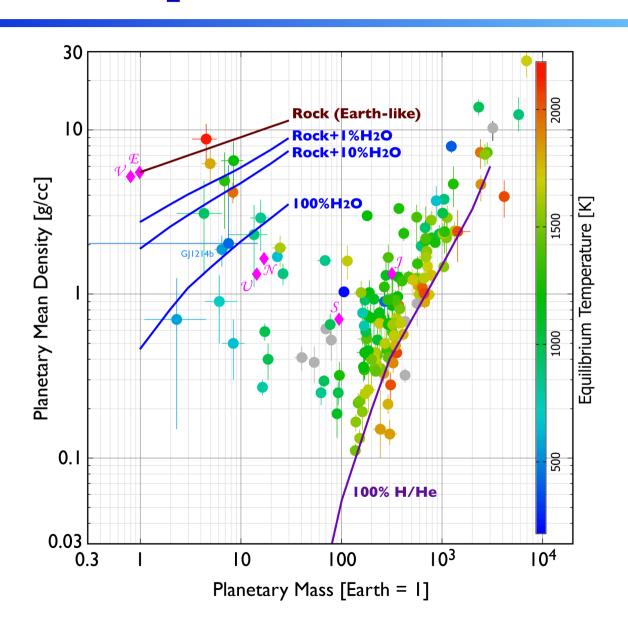
cf. Possibly 5 times ocean mass of water in the Earth's interior



Uranus and Neptune are made mainly of H₂O (60-70%).

A tiny fraction of liquid water on the Earth is essential for origin and evolution of life.

Super-Earths



How Old is the Ocean?

► Isua rocks in West Greenland (3.8 Gyr)

sedimentary rock



pillow-lava



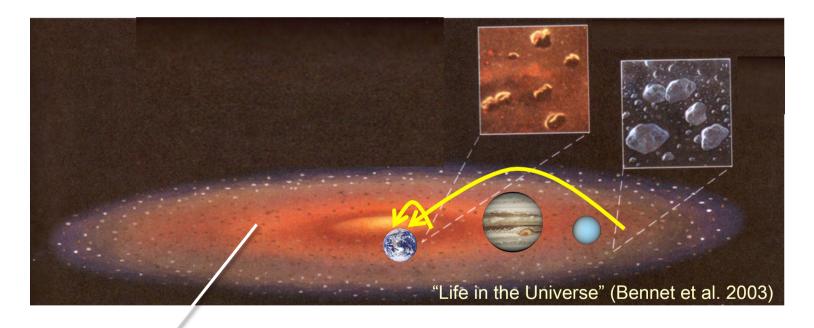
Oxygen isotope of zircons (< ~4.3Gyr) Mojzsis et al. (2001)

These evidences show that the ocean already existed just after (or during) Earth's formation.

Possible Water Sources

(2) Asteroids, Icy planetesimals

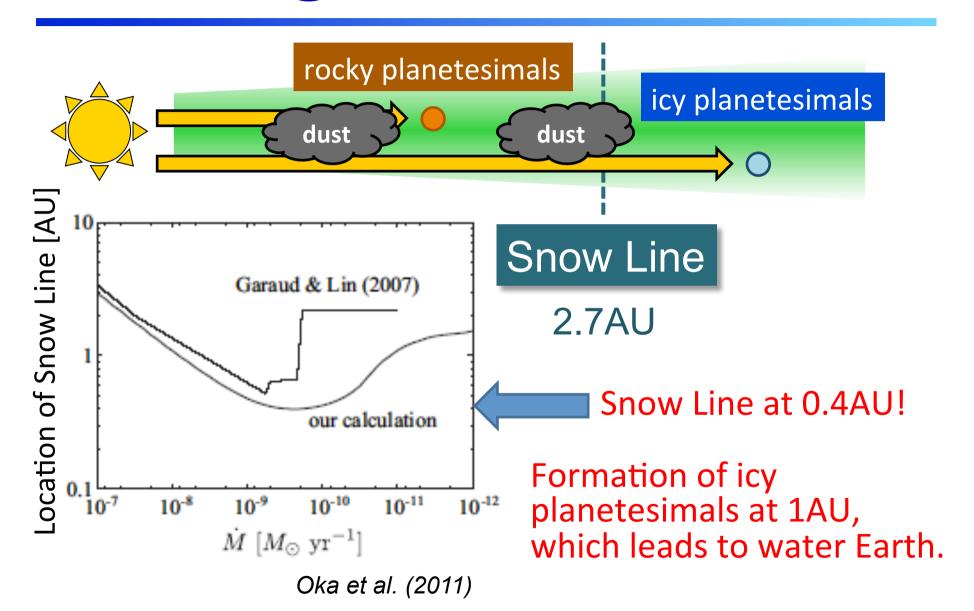
(e.g., carbonaceous chondrites, comets)



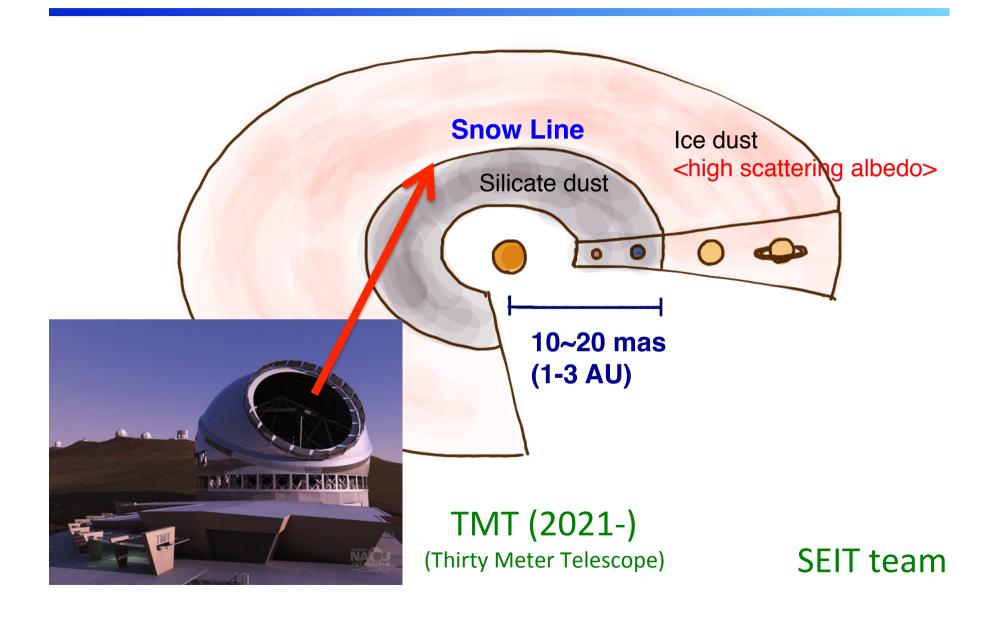
(3) nebular gas

(1) Building block of the Earth (planetesimals)

Building Blocks of Earth

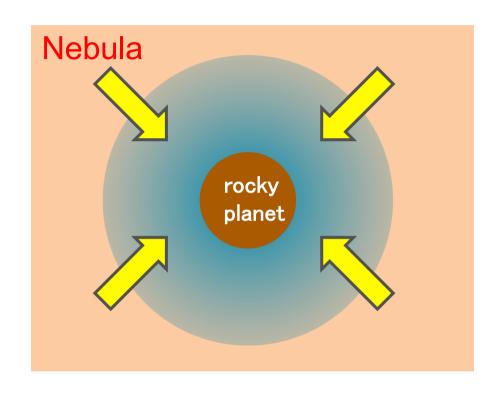


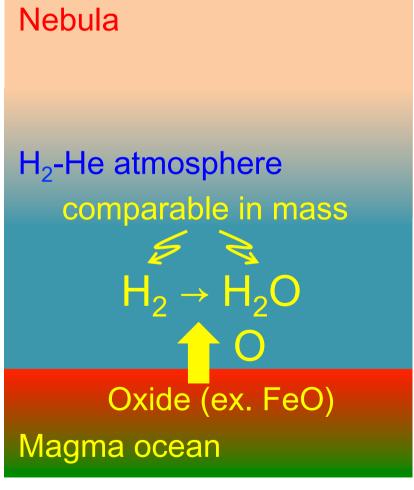
Obs. of Snow Line



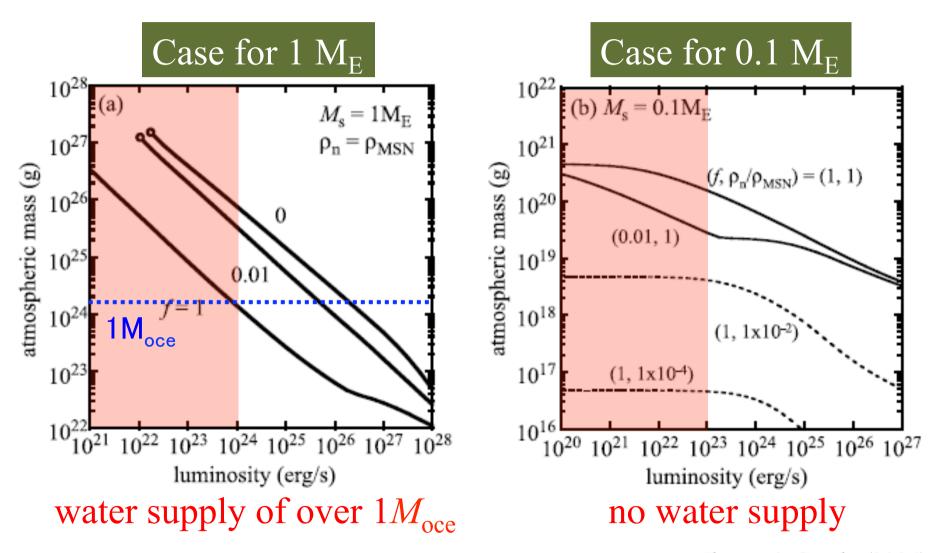
Nebular Origin of Water

Sasaki (1990), Ikoma & Genda (2006)



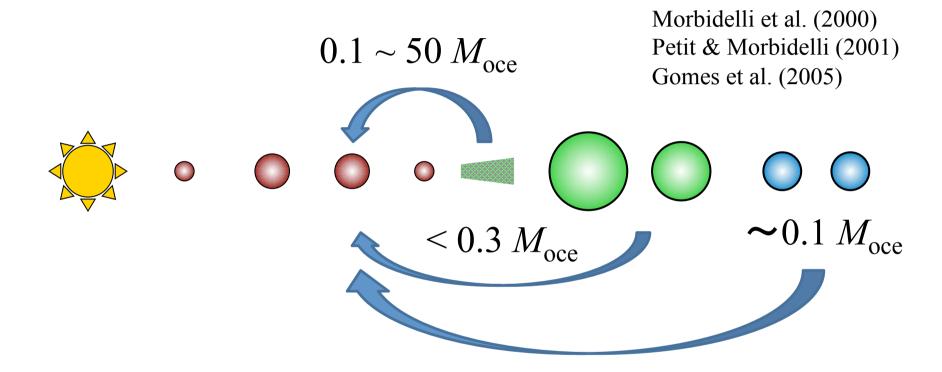


Captured Atmosphere



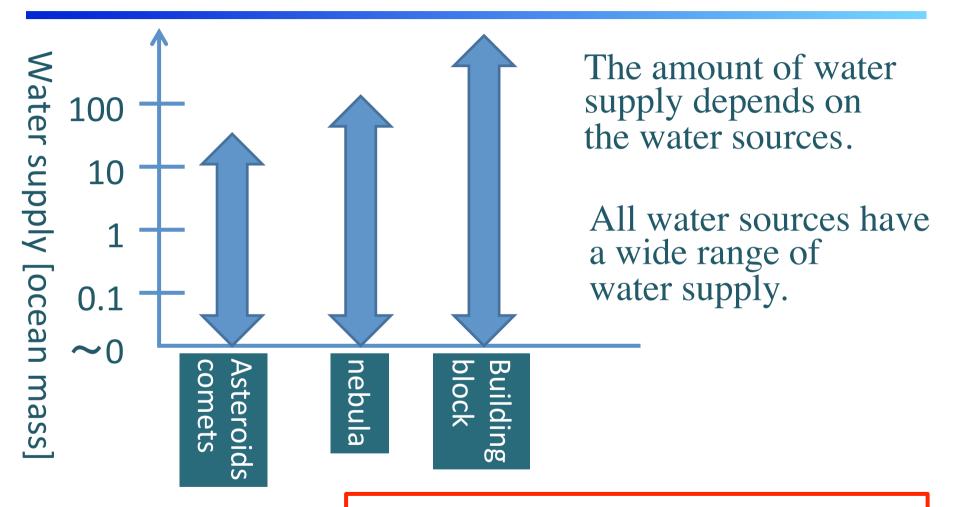
Ikoma & Genda (2006)

Asteroids and Comets



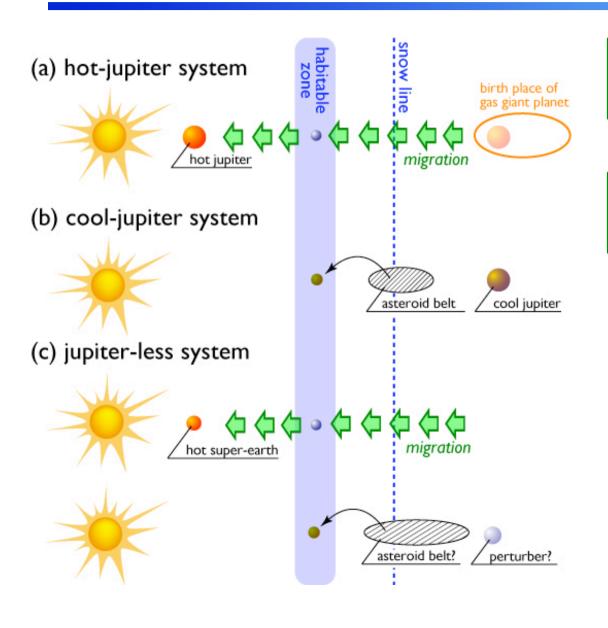
- Contribution of comets is low (blow $0.3M_{\rm oce}$).
- Contribution of asteroids is sometimes huge.

Water Supply



A wide variety of water amount on extrasolar terrestrial planets

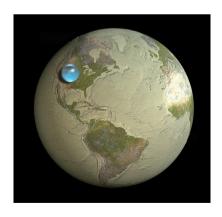
What controls water amount?



wide variety in extrasolar planetary systems



Which type is the best for little water supply?

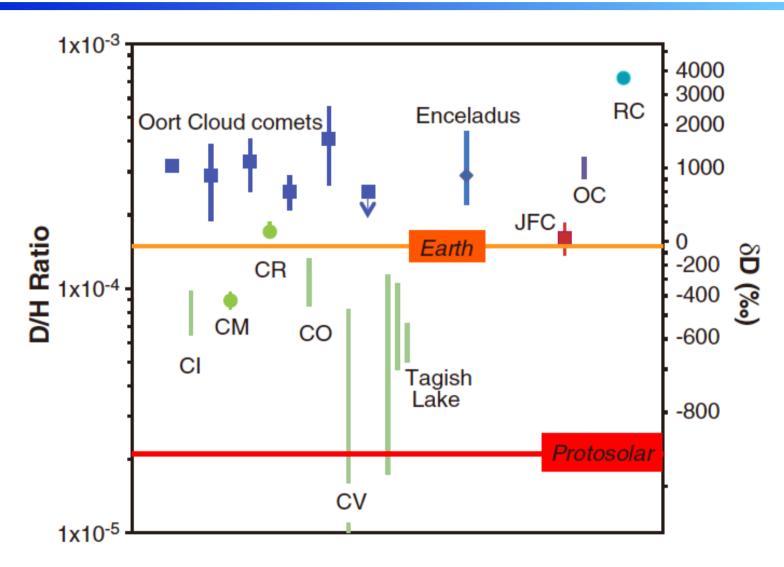


from Maruyama, Ikoma, Genda et al. (2013)

Summary

- A tiny amount of water on the Earth is essential for emergence and evolution of life.
- Source of water on the Earth is under debate.
 - --- Detail investigation of Earth's geology and observation of disks and extrasolar planets will solve this question.
- The theory of planet formation predicts a wide variety of water amount on extrasolar terrestrial planets.

D/H in water sources

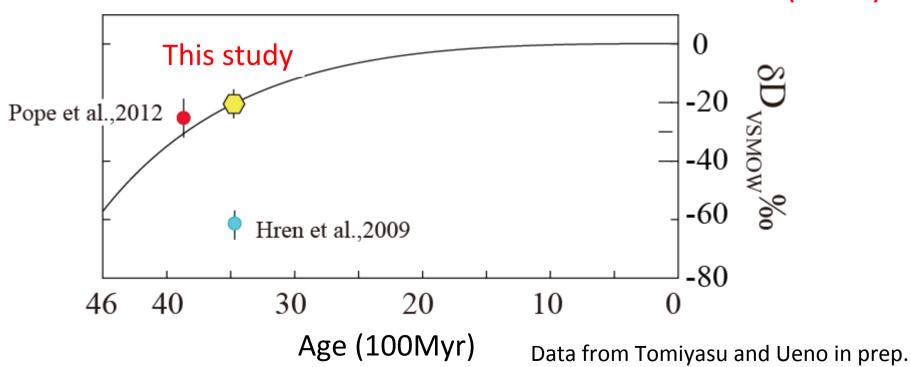


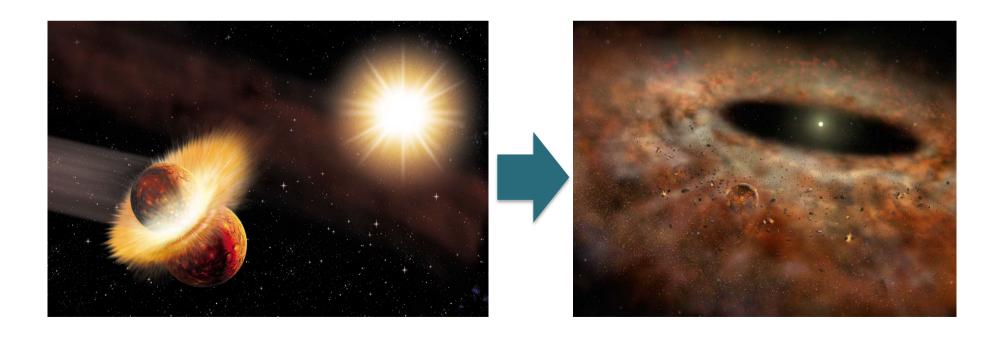
Change of ocean D/H

Previous discussion of D/H is based on the assumption that the D/H of ocean has never varied during 4.5 Gyr.

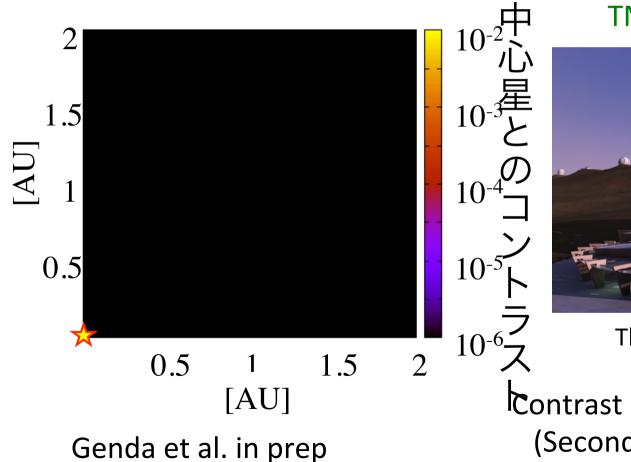
D/H of ocean increases for H₂-rich proto-atmosphere

Genda & Ikoma (2008)

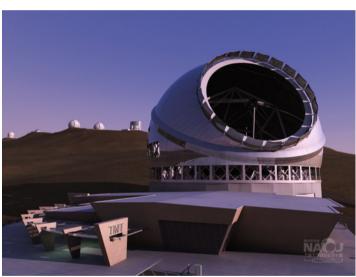




次世代望遠鏡による観測



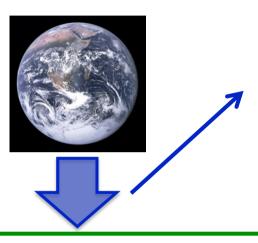
TMT(2021年~)



Thirty Meter Telescope 30m望遠鏡

Contrast: ~ 10⁻⁸ for SEIT (Second Earth Imager for TMT)

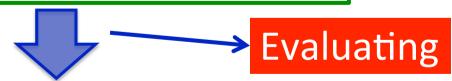
Conditions for Habitable Planet



Extracting

essential conditions for habitable planets from Earth's history

- Liquid water
- Amount of water
- Planetary size etc.



Universality and uniqueness of Earth

these conditions from planet formation theory