1st ELSI International Symposium 2013. 3. 28.

Laboratory Simulation of Formation and Alteration of High Molecular Weight Organics in Space

Kensei Kobayashi Yokohama National University *kkensei@ynu.ac.jp*

Outline

- Classical Scenario of Chemical Evolution
- Problems in the Classical Scenario
- Extraterrestrial Organics and Origins of Life
- Abiotic Formation of Complex (High Molecular Weight) Organics in Space
- Merits of Complex Amino Acid Precursors
- New Scenario of Chemical Evolution
- Genesis & Evolution of Biochemical Functions

Classical Sinario of Chemical Evolution

- Miller (1953): Amino acid synthesis from strongly reducing atmosphere
- Oro (1960):

Adenine synthesis from HCN solution

• Orgel et al. (1970s-80s): Abiotic syntheses of oligonucleotides





Problems in Classical Scenario (1) High concentration (~1M) of pure starting meterials are required

- HCN + HCHO + NH_3
 - \rightarrow Amino acids (Strecker Reaction)
- 5 HCN \rightarrow Adenine
- 5 HCHO \rightarrow Ribose (Formose Reaction)
- Inhibitors should not co-exist (*eg.* Ammonia in Formose reactions)
- Amino acids and sugars are easily decomposed in aqueous solution by heat or radiation

Problems in Classical Scenario (2) Regio- and Stereoselectivity



Correct and Wrong Bonds

 α -pyranoside, β -pyranoside....

Problems in Classical Scenario (3) Primitive Earth Atmosphere was not Strongly-Reducing



- Primitive atmosphere was only slightlyreducing
- Amino acids are hardly formed from slightly-reducing gas mixtures by UV, heat and spark discharges

http://www.newtonpress.co.jp/search2/e_book/ebook/2002_02/mmsolarsys_0502.html

New Insight 1: Reevaluated Energies for Prebiotic Syntheses

 $\begin{tabular}{l} \hline \underline{Materials} & : \\ Weakly-reducing gas mixtures \\ (CO_2, CO, N_2, H_2O) \\ (N_2, CH_4 (tr.), H_2O) \\ \hline \end{tabular}$

Energies:

Cosmic rays (Proton irradiation) Bolide impacts (MPD arc jet)

Amino acid precursors with high molecular weights are produced





New Insight 2: Complex Organics in Space



New Insight 3: Seeds of homochirality in Space

соон

NH

R

Mirror Image

Enantiomeric excesses Found in meteorites

Murchison meteorite

H

Extended Circularly-polarized light in the Orion massive star forming region (Fukue et al., 2011)

соон

R

NH

9

太陽系の約400倍

太陽系の約100倍

Origin of Organics in Comets and Meteorites: Formation in Molecular Clouds?



Molecular cloud

Greenberg Model:

Complex organics are formed in ice mantles of interstellar dust particles in molecular clouds

Irradiation of Simulated ISD Ices

Charged Particles

- Kobayashi et al., 1995
- Kasamatsu et al., 1997
- Kobayashi et al., 2007

Ultraviolet Light

- Munos Caro et al., 2002
- Bernstein et al., 2002
- Takano et al., 2003



Particles Irradiation of Simulated Interstellar Media



Heavy ions@HIMAC(NIRS) \downarrow





- Molecular weights: Some thousands
- Amino acids were yielded after hydrolysis
- ✓ Catalytic activities

Merits of Complex Amino Acid Precursors (1) Stability of Free Glycine and Complex Glycine Precursors* against v-ravs irradiation



•CAW: Product when a mixture of CO, NH3 and H2O was irradiated with High Energy Protons.

Merits of Complex Amino Acid Precursors (2)Origin of Bio-homochilarity

L', D': Amino acid precursors

Enantiomeric excesses formed without assymetric destruction



Merits of Complex Amino Acid Precursors (3)boundaries



New Senario of Origins of Life: The Garakuta * World



Features of LIFE



Evolution of Biochemical Features



We will be able to Find Garakuta Molecules and Garakuta Life in:





Thank you for your attention!

Acknowledgement

- Dr. Satoshi Yoshida & HIMAC Staff (NIRS)
- Dr. Katsunori Kawasaki (TIT Van de Graaff)
- Dr. Hitoshi Fukuda & Dr. Yoshiyuki Oguri (TIT Tandem)
- Dr. Kazuhiro Kanda & NewSUBARU Staff (Univ. Hyogo)
- Dr. Jun-ichi Takahashi (NTT)
- Dr. Yoshinori Takano (JAMSTEC)
- Members of Kobayashi Lab(YNU)